

PhotoDet 2015

International Conference on New Photo-detectors
July 6-9, 2015, Moscow, Troitsk, Russia



Study of scintillation counter consisting of a pure CsI crystal and APD

D. Epifanov (The University of Tokyo)

on behalf of Belle II collaboration

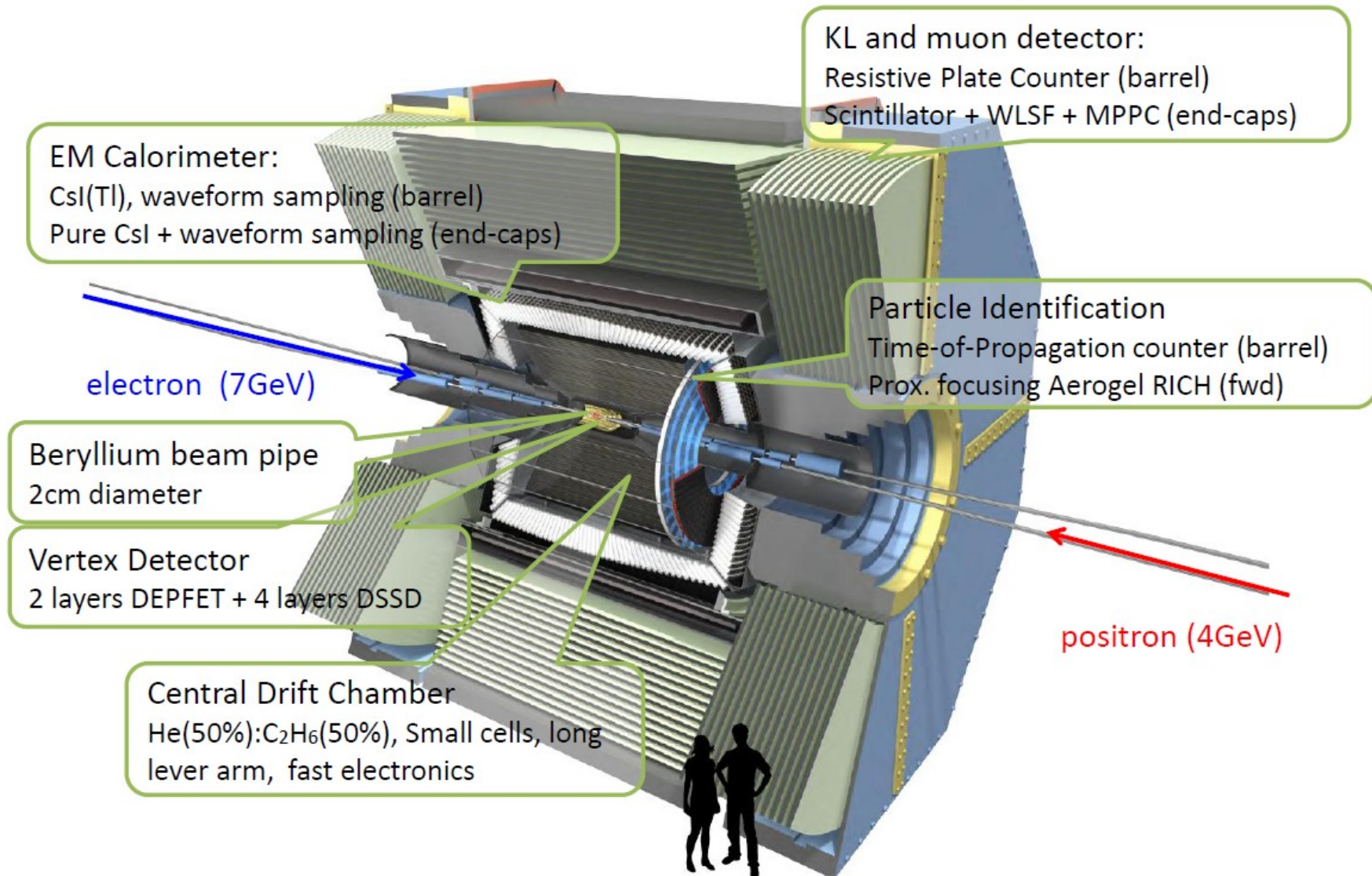
July 9th, 2015

Outline:

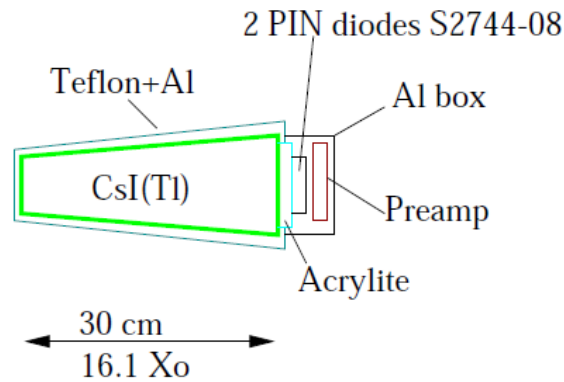
- Belle II calorimeter upgrade
- Electronic noise in the scheme with APD
- CsI(pure)+(1-4)APDs light output and equivalent noise energy
- Improvement of the light output
- Wavelength shifters with organosilicon luminophores
- Characteristics of APD
- Summary

Belle II at SuperKEKB

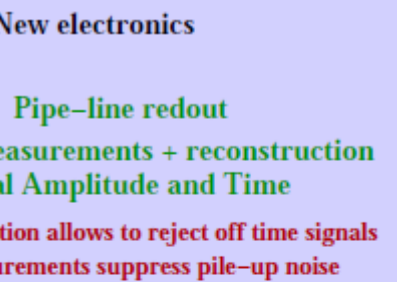
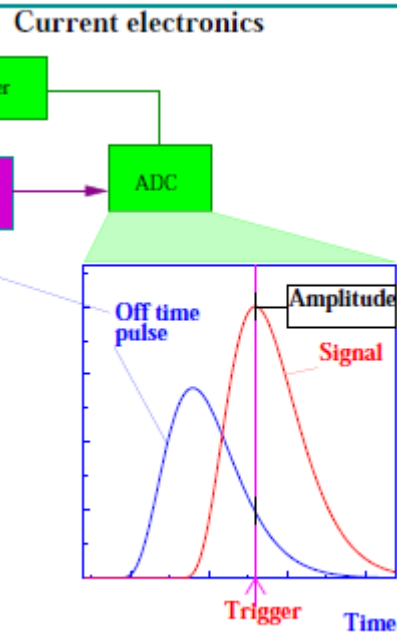
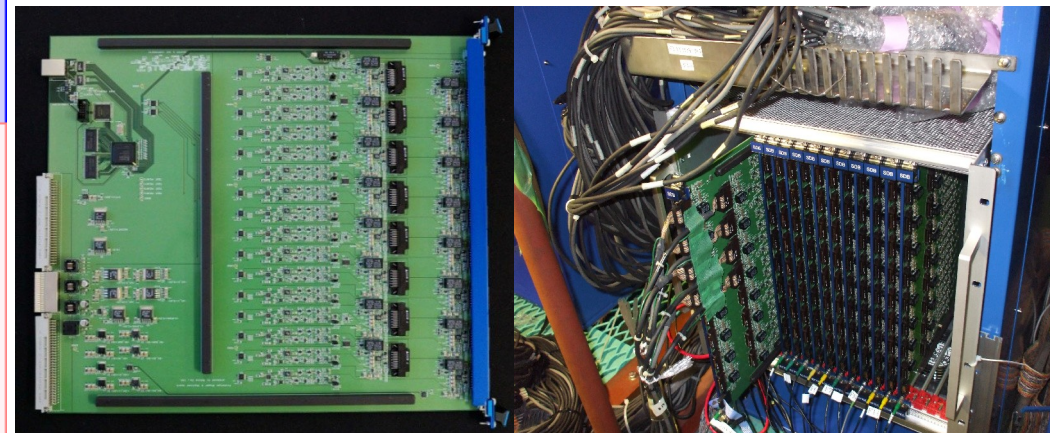
Belle II at SuperKEKB is the only e^+e^- Super Flavor Factory in the nearest future, which is competitive/complementary to the current and coming energy/intensity frontier experiments



Barrel ECL

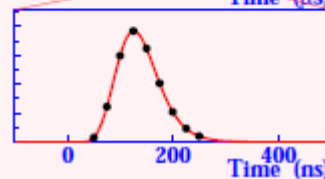
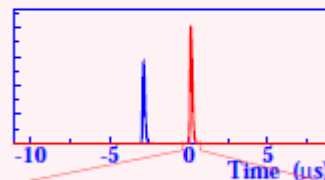


- **Barrel ECL** will be reused, new electronics with pipe-line readout and waveform analysis (16 ch Shaper-DSP board) has been developed and tested. All Shaper-DSP boards were produced, tested, delivered to KEK and installed in the detector.
- Belle II DAQ electronics has been tested in the ECL data transfer runs with the frequency up to 30 kHz.
- **Cosmic runs with barrel ECL are ongoing.**



Pipe-line readout
 ■ 16 measurements + reconstruction
 Signal Amplitude and Time

- Time information allows to reject off time signals
- Several measurements suppress pile-up noise



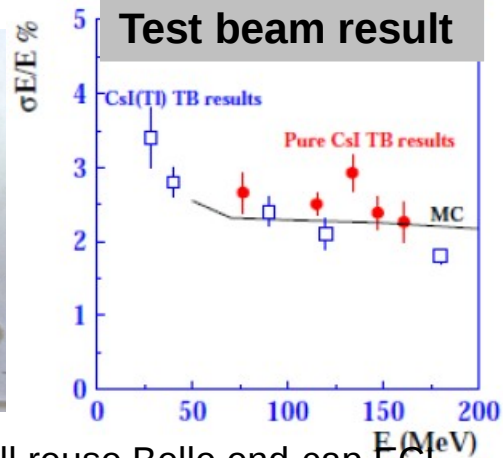
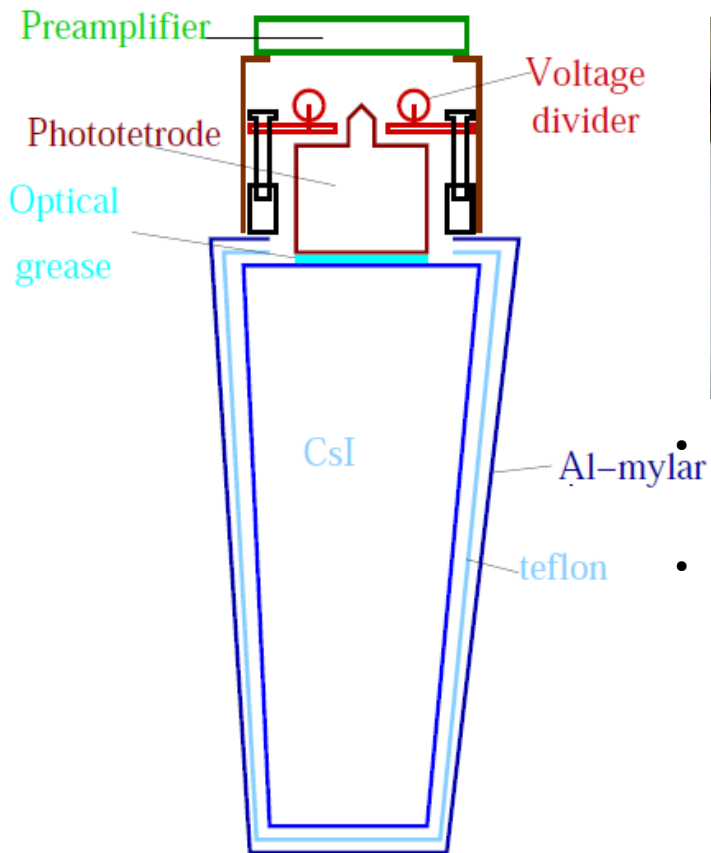
Pure CsI for endcaps
 CsI(Tl) $\tau=1\mu s$
 PIN diodes

→

pure CsI
 $\tau=30 ns$
 Vacuum phototetrodes

Essentially better time resolution ($\sigma=1 ns$)
 Essential pile-up noise suppression

End cap ECL upgrade

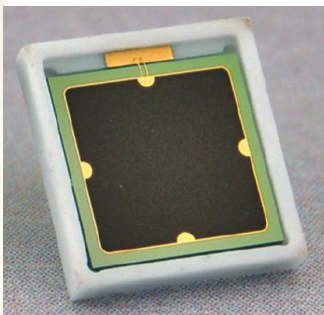


At the first stage of the Belle II experiment we will reuse Belle end-cap ECL (1152 + 960 channels) (with new preamplifiers and readout electronics).

- **The main end-cap ECL upgrade option is to use CsI(pure) crystals and Hamamatsu photopentodes R11283MOD-A** (dedicated R&D showed good results):
 - Low pile-up noise and good energy and spatial resolution
 - Similar physical characteristics (as for CsI(Tl)), better radiation hardness
 - There are several crystal producers, acceptable price

Hamamatsu APD S8664-1010

EXCELITAS APD C30739ECERH-2



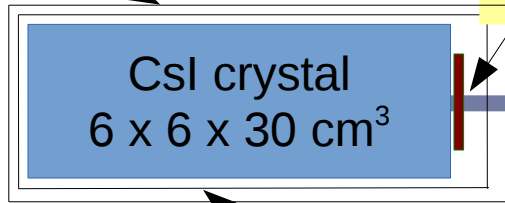
• **However there are some difficulties:** *no redundancy, notable dependence on magnetic field, long term stability, new mechanical support is needed.*

• **In the CsI(pure) + Si APD option** we are investigating APDs from two producers: *EXCELITAS, Hamamatsu Photonics.*
The main problem here is to reach admissible level of electronic noise.

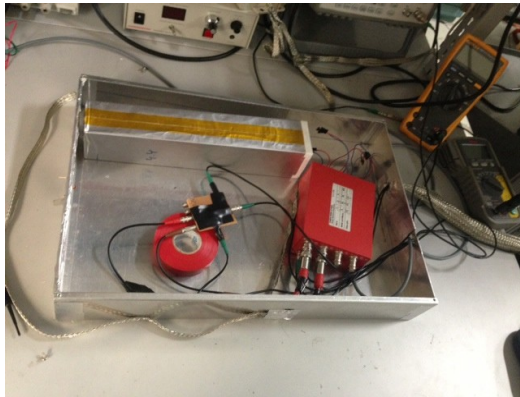
• **With the actual size crystal and 1 APD (1 x 1 cm²) Hamamatsu S8664-1010 we obtained ENE ≈ 2 MeV, while the required ENE ≤ 0.5 MeV**

Study of electronic noise

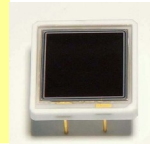
Aluminized mylar



Teflon



APD Hamamatsu S8664-1010
 $U_{\text{bias}} = 394 \text{ V}$: Gain = 50, $I_{\text{dark}} = 8 \text{ nA}$ @ $T=25 \text{ }^\circ\text{C}$



CAEN preamp.

Shaper

ADC

PC



4ch preamplifier
CAEN A1422B045F3
 45 mV/MeV (1 V/pC)

CP 4467A
 Fast Shaping
 Amplifier (NIM)
 $\tau = (20\text{—}500) \text{ ns}$

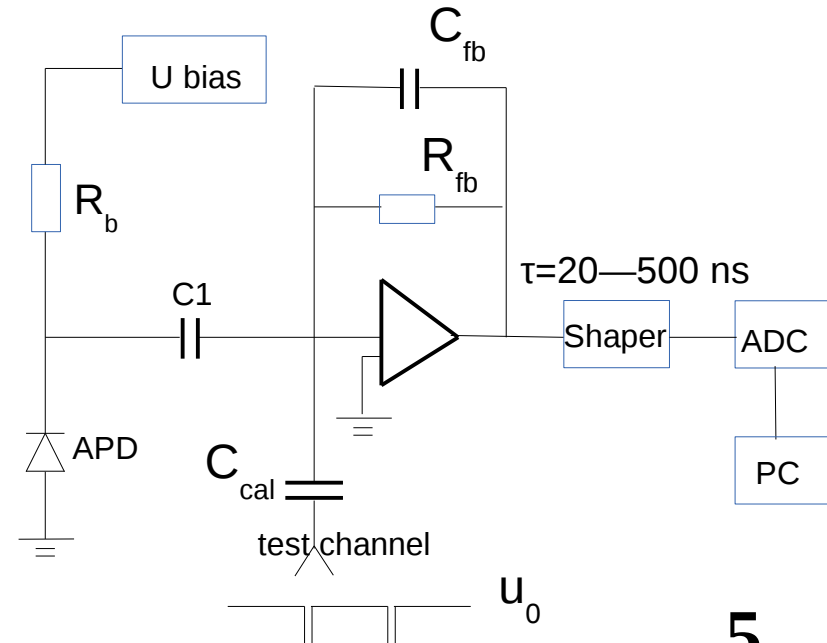


Hoshin C008
 16ch peak hold
 ADC (CAMAC)

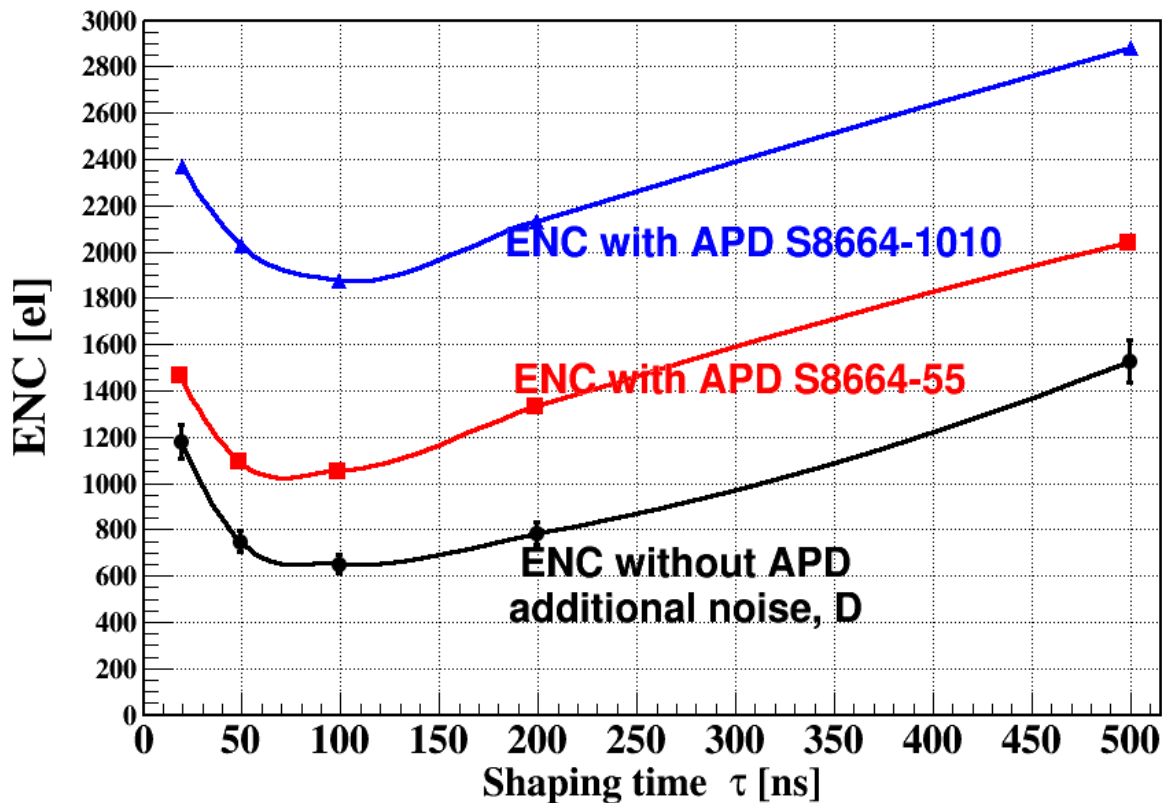
$$ENC^2 = \frac{2I_d K g F \tau}{e} + \left(\frac{B^2}{\tau} + E^2 \right) C^2 + D^2$$

Shot noise
Thermal noise
Additional noise

- | | |
|--------------------------------------|---------------------------------------|
| e – positron charge; | F – excess noise factor; |
| I_d – dark current; | C – APD junction capacitance; |
| g – APD gain; | B – thermal noise coefficient; |
| τ – shaping time; | E – 1/f noise coefficient; |
| K – shaper factor; | D – additional noise. |



Measurement of D



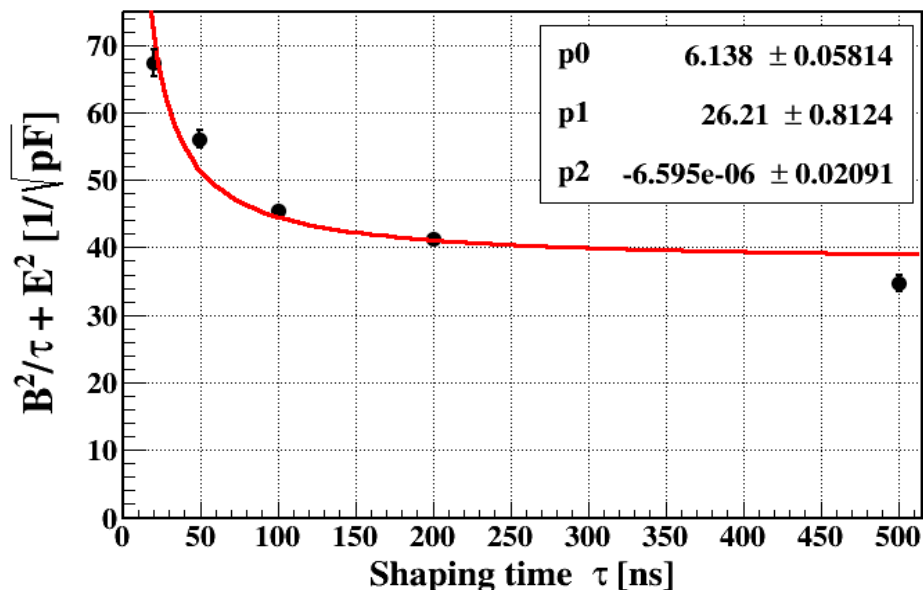
- At the shaping times from 20 ns to 500 ns, D is not constant. It varies strongly, which is explained by the relatively large additional parallel (i_{na}) and serial (e_{na}) noises.
- Fast shaper of better quality (like ORTEC 474, 579) might be helpful to decrease D

$$ENC^2 = (2eI_d + \frac{4k_bT}{R_b} + i_{na}^2) K_i T_s + (4k_bT R_s + e_{na}^2) K_v \frac{C^2}{T_s} + K_{vf} A_f C^2$$

Measurement of thermal noise (B, E)

Two well known capacitors C_1 and C_2 were used to measure **B** and **E**.

$$\mathbf{B}^2/\tau + \mathbf{E}^2 = (\overline{Q_1^2} - \overline{Q_2^2}) / (C_1^2 - C_2^2)$$



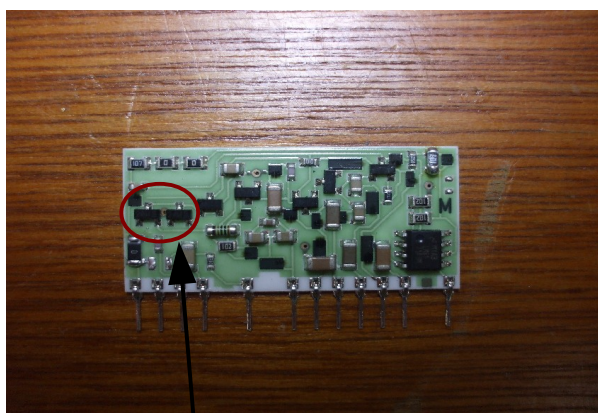
$$\mathbf{B} = (26.2 \pm 0.8 \pm 4.8) \sqrt{\text{ns/pF}}$$

$$\mathbf{E} = (6.1 \pm 0.1 \pm 0.4) \text{ 1/pF}$$

$$\mathbf{B}^2/\tau = (4k_B T R_s \Delta f) / e^2$$

R_s , equivalent serial resistance, it is dominated by reversal transconductance of the CAEN preamp. FET (BF862)

$R_s \approx 50 \Omega$ was also measured with additional serial resistance at the CAEN preamp. input



2 BF862 FETs

We also tried FET **2SK932-23**, at short shaping times thermal noise is almost the same.

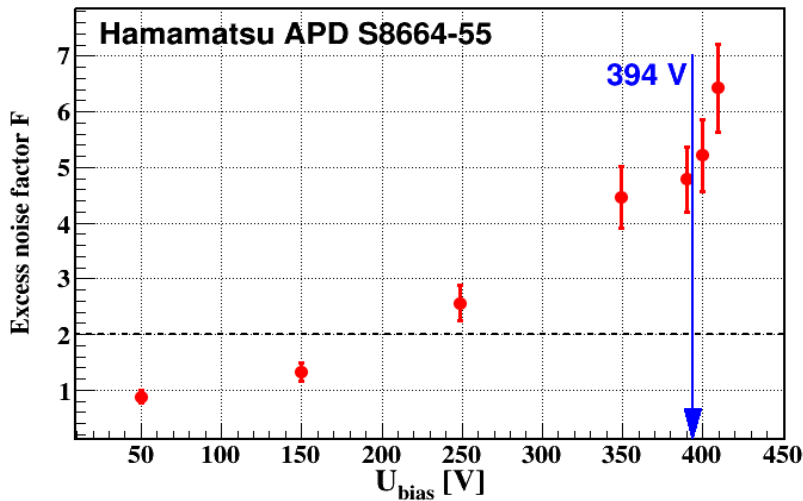
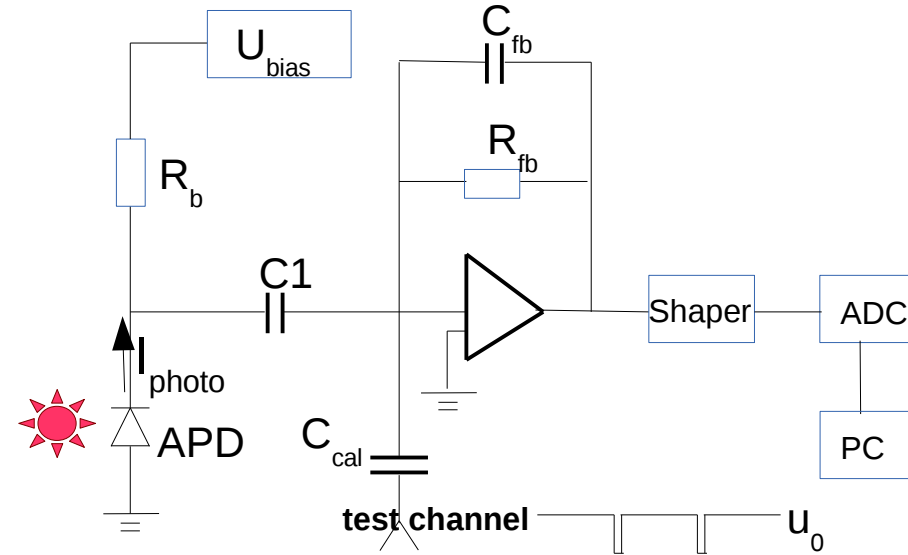
Shot noise, excess noise factor F

$$\overline{Q^2}_{\text{no } I_{\text{photo}}} = 2 \cdot e \cdot I_d \cdot \tau \cdot g \cdot F \cdot K + (B^2/\tau + E) \cdot C_d^2 + D^2$$

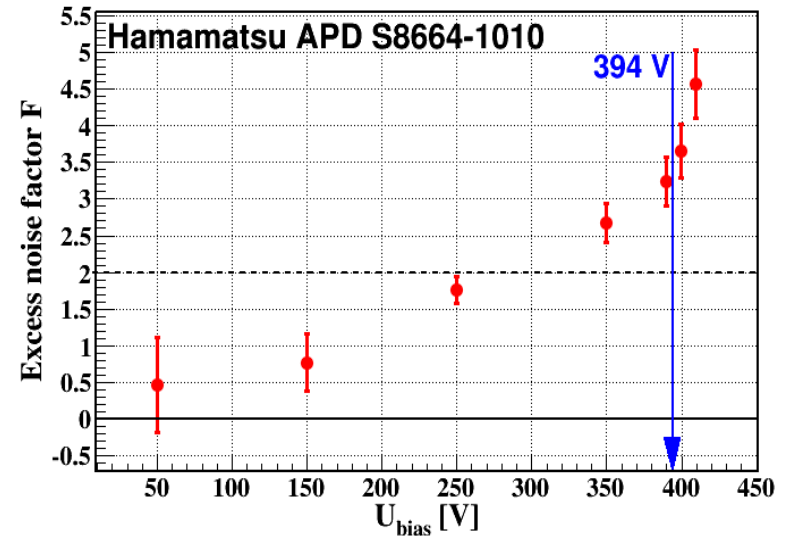
$$\overline{Q^2}_{\text{with } I_{\text{photo}}} = 2 \cdot e \cdot (I_d + I_{\text{photo}}) \cdot \tau \cdot g \cdot F \cdot K + (B^2/\tau + E) \cdot C_d^2 + D^2$$

$$F = (\overline{Q^2}_{\text{with } I_{\text{photo}}} - \overline{Q^2}_{\text{no } I_{\text{photo}}}) / (2 \cdot e \cdot I_{\text{photo}} \cdot \tau \cdot g \cdot K)$$

$$K(\text{EXP}) = 0.44 \pm 0.02 \quad K(\text{CR-4RC}) = 0.45$$

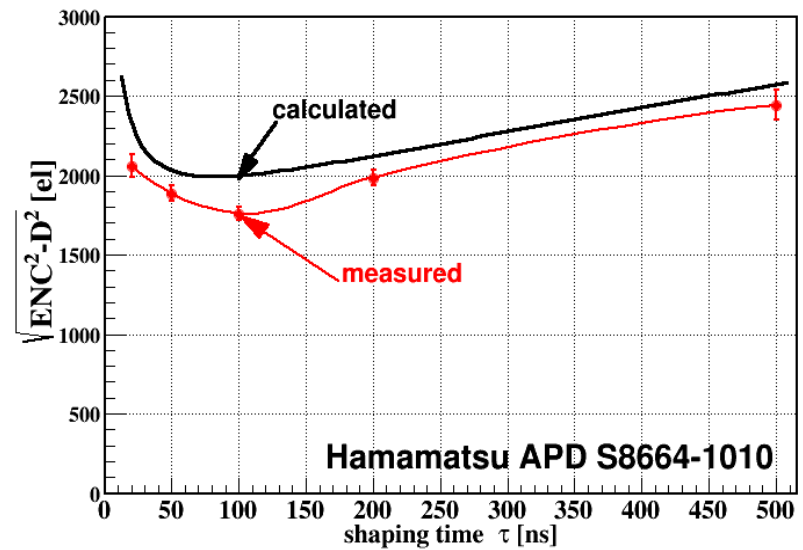
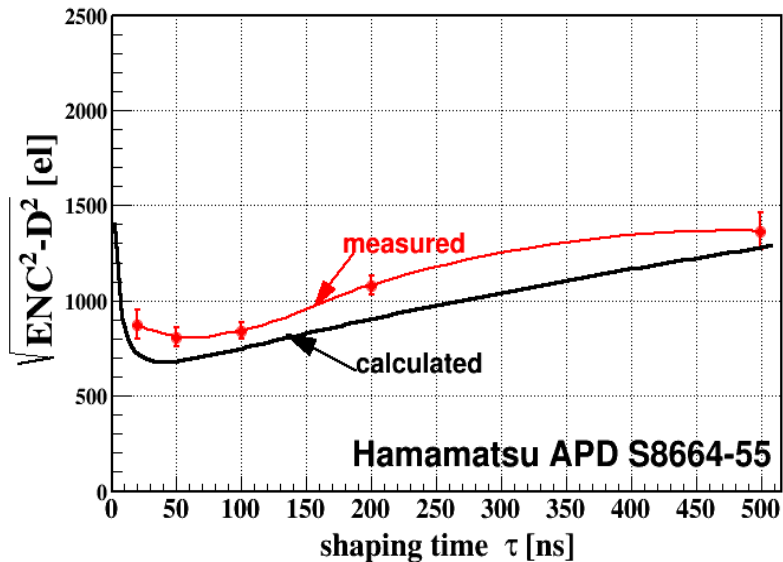
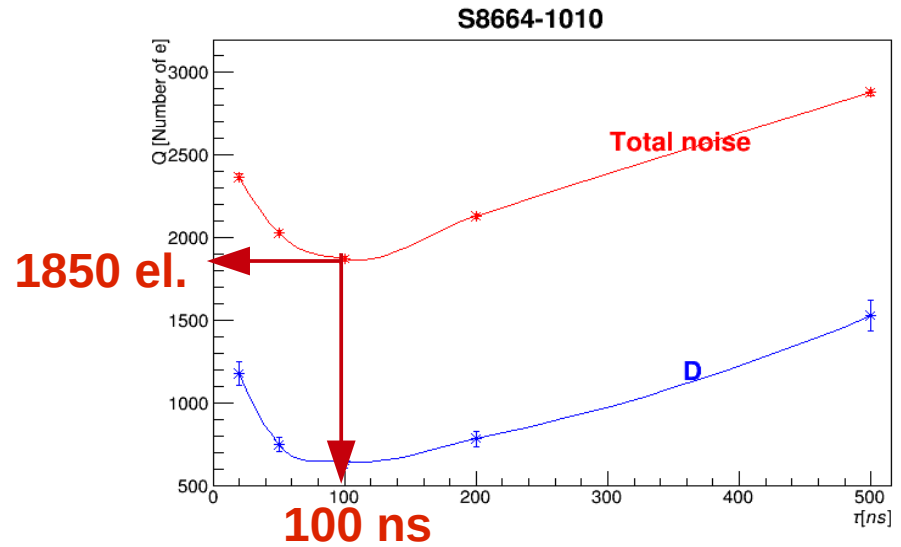
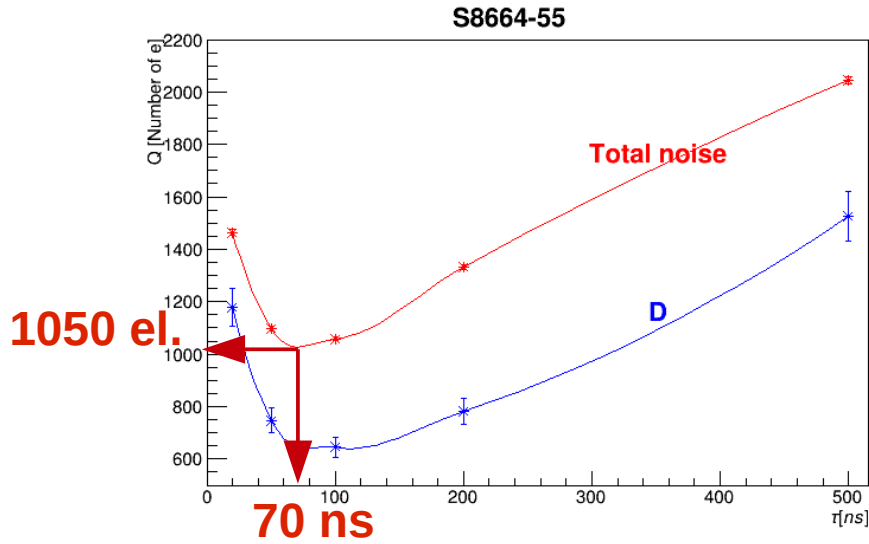


S8664-55: $g = 50$, $F = 5.1 \pm 0.5$



S8664-1010: $g = 50$, $F = 3.4 \pm 0.4$

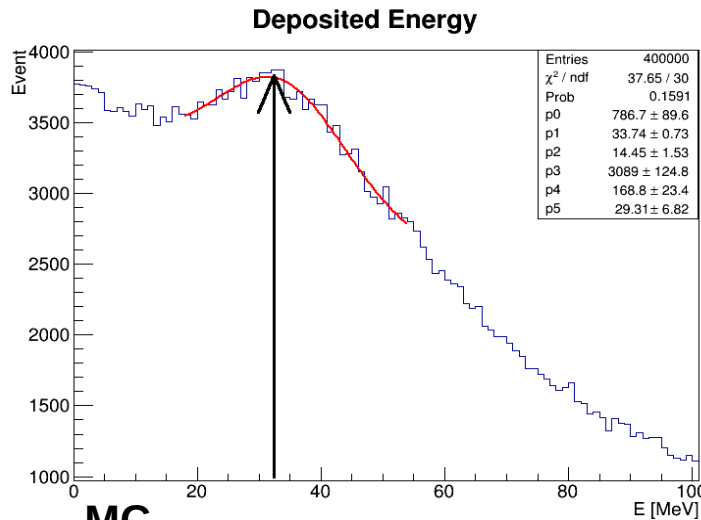
ENC vs. shaping time



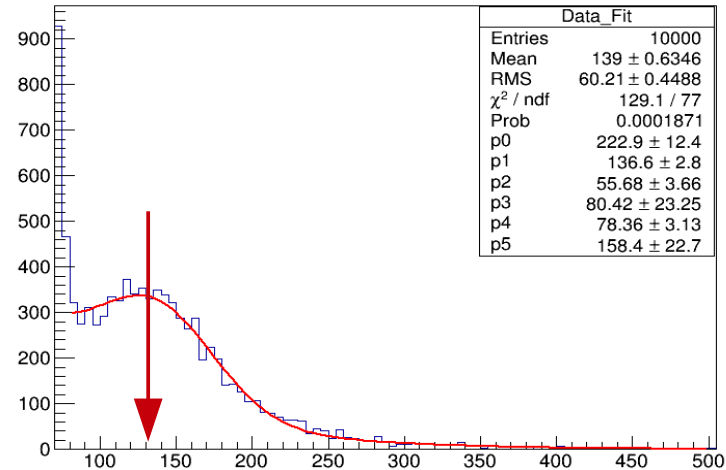
Discrepancy between **calculated** and **measured** $\sqrt{\text{ENC}^2 - \text{D}^2}$ is due to the uncertainty in \mathbf{C}_{APD}

Light output (LO) and ENE

Cosmic muons are used to calibrate ADC channels in units of energy (MeV)



$$E_{\text{peak}}^{\text{MC}} (\text{cosmic}) \approx 33 \text{ MeV}$$



$$\text{Price}_{\text{ADC}} (\text{MeV/ch}) = E_{\text{peak}}^{\text{MC}} / A_{\text{peak}}^{\text{EXP}}$$

$$\text{ENE} = \sigma_{\text{cal}} \times \text{Price}_{\text{ADC}}$$

The light output is measured by comparison of the signal from cosmic muons (A_{cosm}) with calibration signal (A_{cal}) (gain is eliminated)

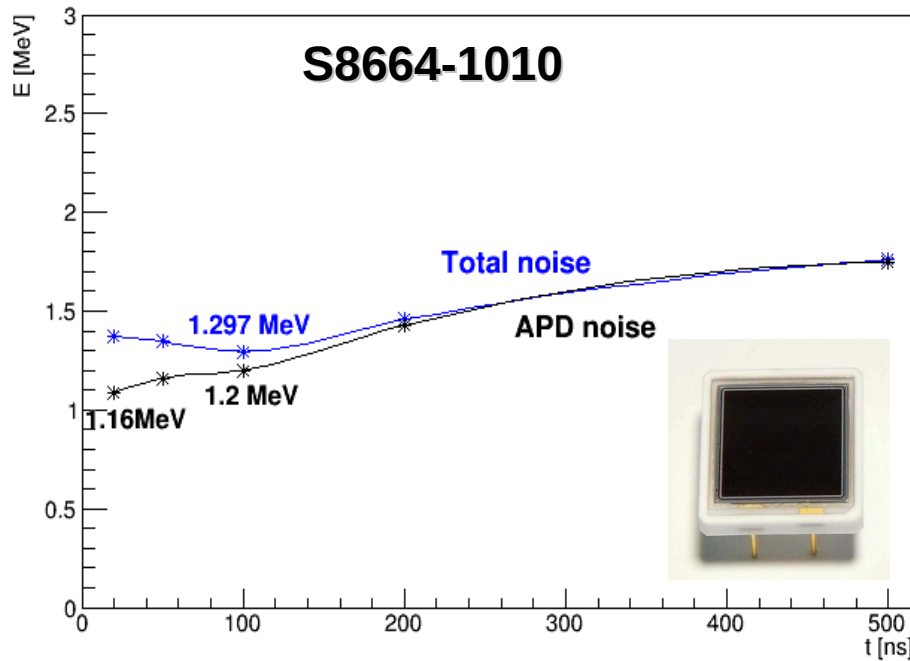
$$N_{\text{cosm}} (\text{ph.e.}) = (C_{\text{cal}} \times U_0 / e) \times (A_{\text{cosm}} / A_{\text{cal}})$$

$$\text{LO} = N_{\text{cosm}} / E_{\text{peak}}^{\text{MC}} / (\text{APD gain} = 50) / (S_{\text{APD}} [\text{cm}^2])$$

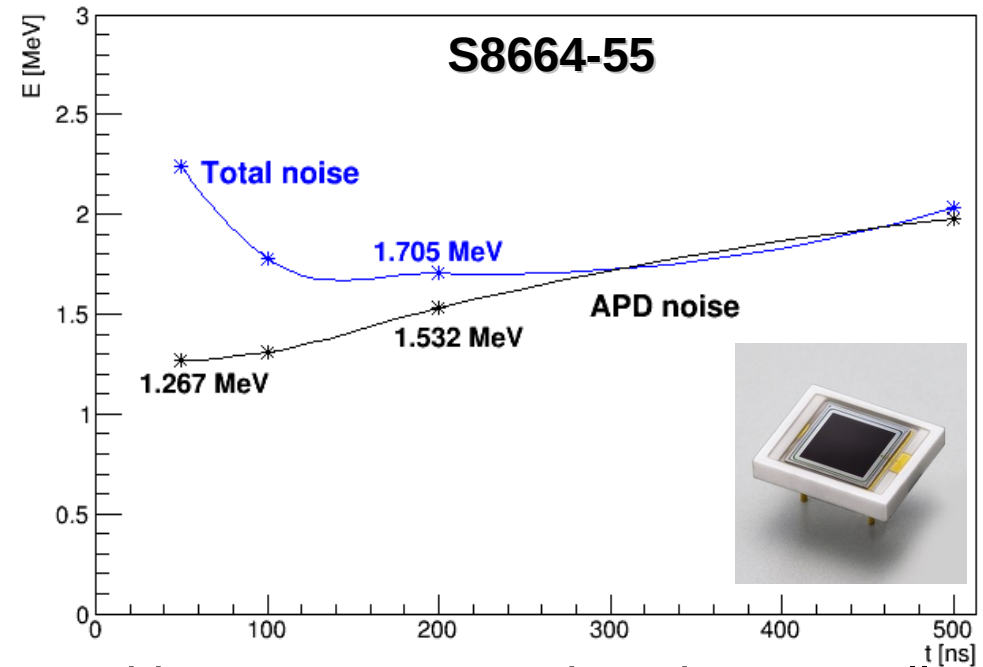
$$\text{LO} = 26 \text{ ph.e. / MeV / cm}^2$$

ENE, several APDs per crystal

2 large APD+6*6*30cm³ CsI Crystal+200um Teflon



2 small APD+6*6*30cm³ CsI Crystal+200um Teflon



Light collection efficiency for the counter with S8664-55 APD is 4 times smaller, than for the counter with S8664-1010, but the thermal noise component is also smaller by a factor of $C_{APD}(\text{large}) / C_{APD}(\text{small}) \approx 3.5$

1 APD S8664-1010 has essentially larger dark current (26 nA) in comparison with the average one (8 nA), we introduce correction to ENE

ENE(2 S8664-1010 APDs (same I_{dark})) = 1.1 MeV

ENE(4 S8664-1010 APDs (same I_{dark})) = 0.8 MeV

ENE(2 S8664-55 APDs) = 1.7 MeV

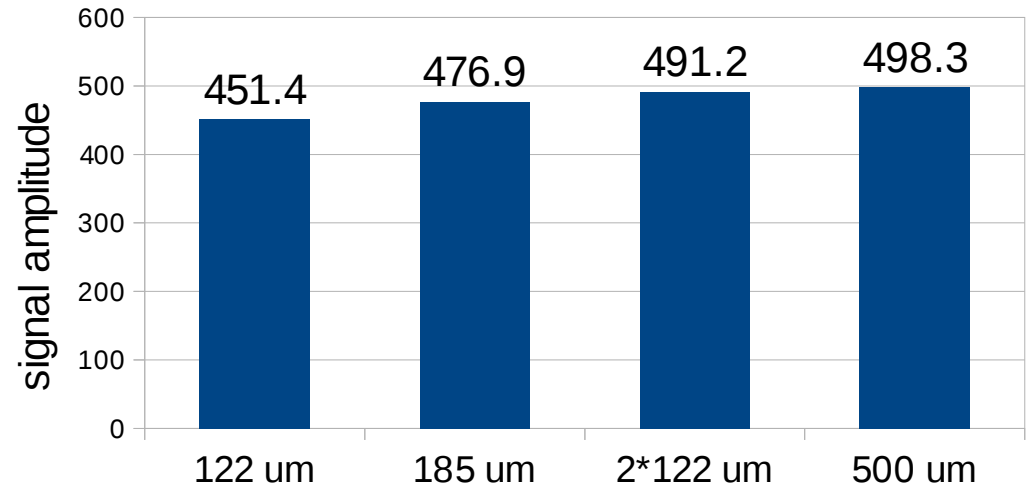
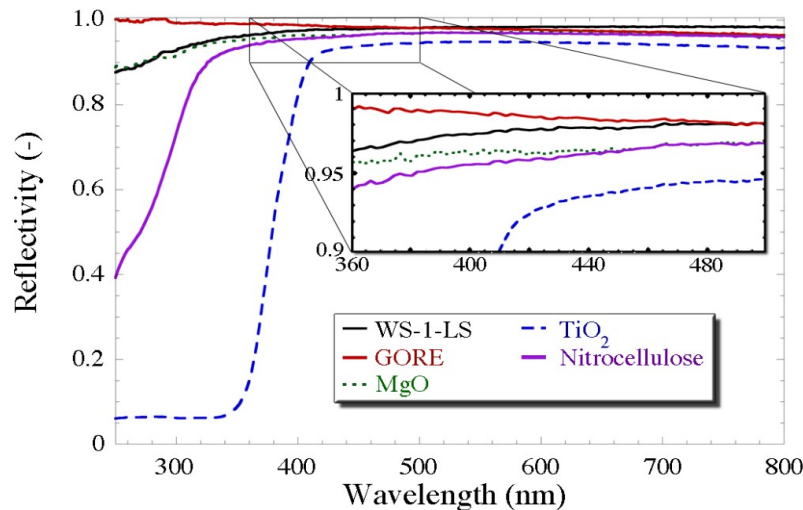
ENE(4 S8664-55 APDs) = 1.2 MeV

Improvement of the LO

	Refraction index	Transparency @315 nm from the producer	Light collection efficiency
OKEN-6262A	1.453 (@ 590 nm)	85%	1.00
TSF451-50M	1.404 (@ 590 nm)	98%	0.85
BC-630	1.465	95%	0.95

Three types of optical grease were tested ($\Delta = 100 \mu\text{m}$),
OKEN-6262A provide the largest light output

Effect of the thickness of white porous Gore-Tex teflon was studied,
 thickness of **200 μm** was found to be optimal.

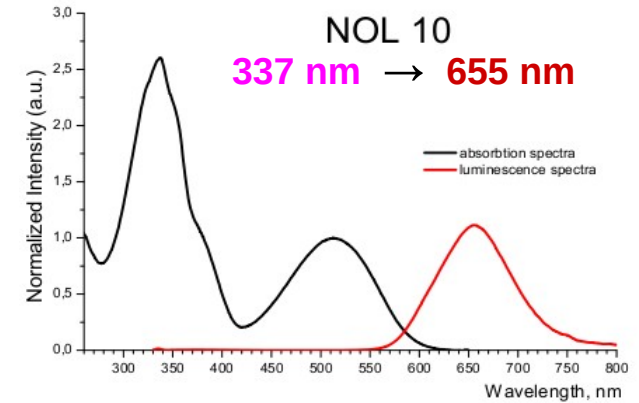
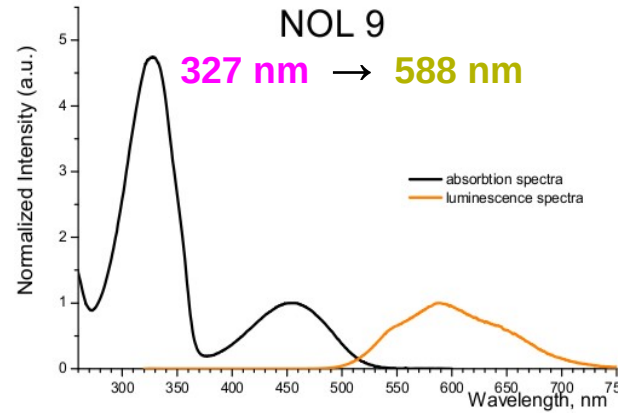
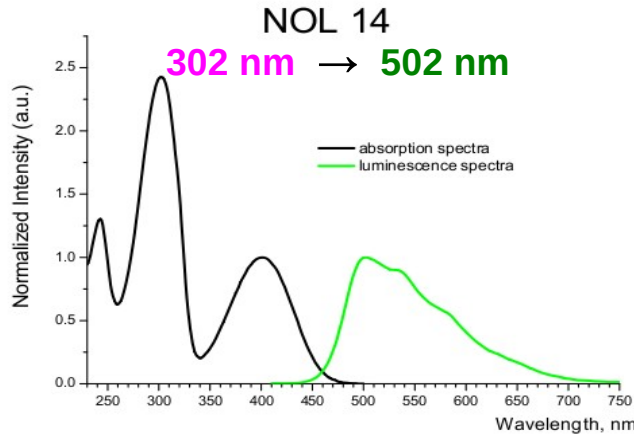


M.Janecek, IEEE Trans. Nucl. Sci. 59.3 (2012) 490.

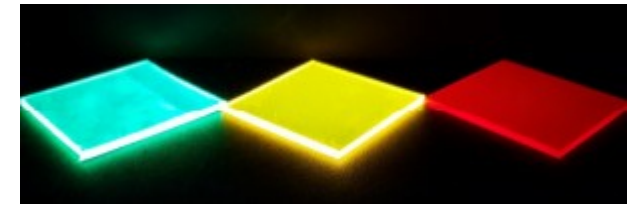
thickness of teflon

Wavelength shifters with organosilicon luminophores

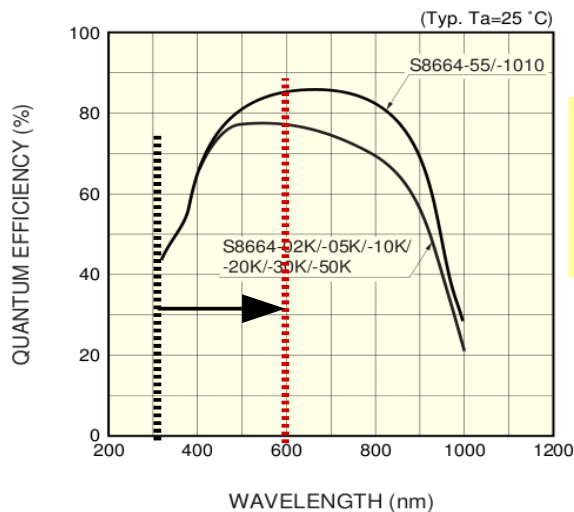
Based on the nanostructured organosilicon luminophores (NOL-9,10,14) from **LumInnoTech Co.**, the WLS plates were developed ((60 x 60 x 2) mm³).



The absorption and emission spectra of these NOL's match our needs very well ($\lambda_{\text{Csi}} = 320 \text{ nm}$). The improvement of the APD QE is by a factor of 2–3.



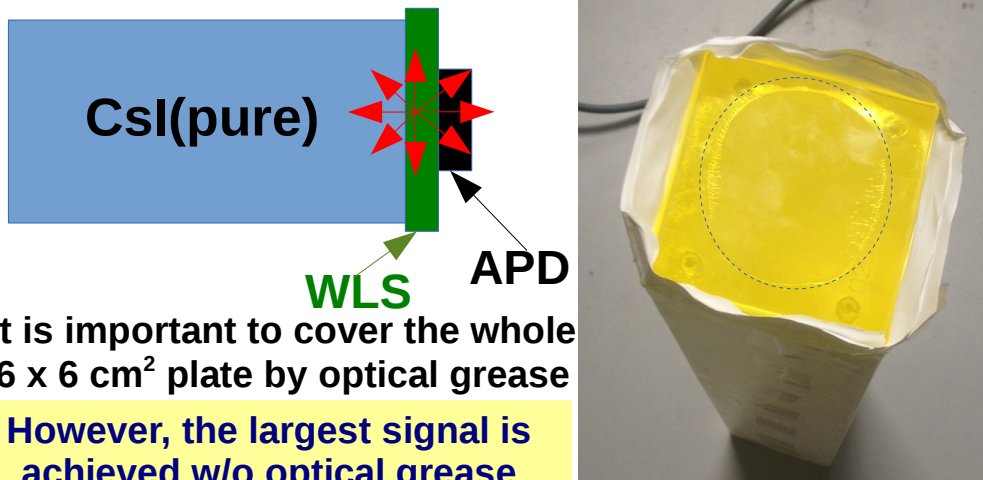
■ Quantum efficiency vs. wavelength



See also poster at this conference:

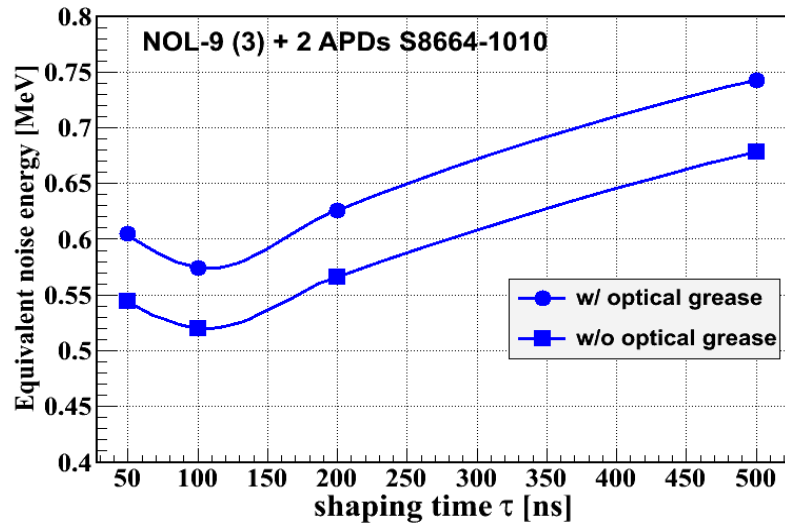
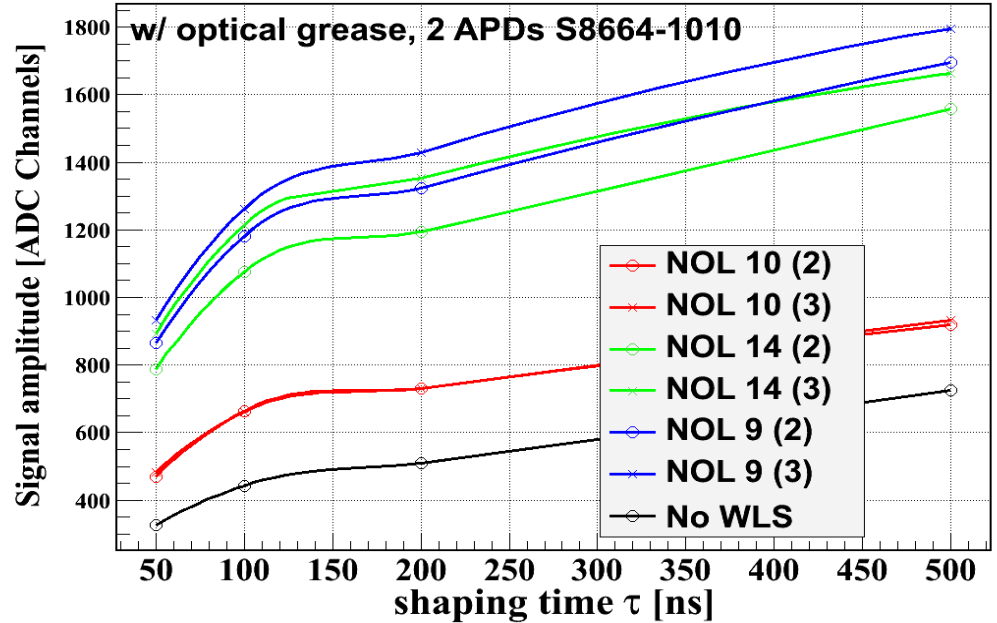
O. V. Borshchev et al., (ISPM RAS)
Nanostructured organosilicon luminophores as effective spectral shifters in a wide spectral region

Results with WLS plates

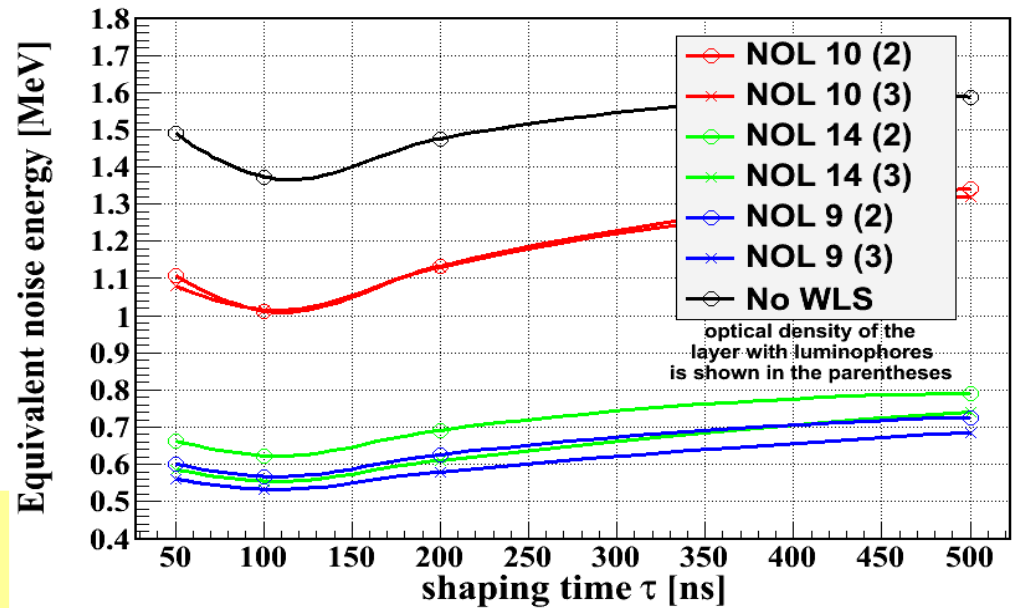


It is important to cover the whole 6 x 6 cm² plate by optical grease

However, the largest signal is achieved w/o optical grease between crystal and WLS plate

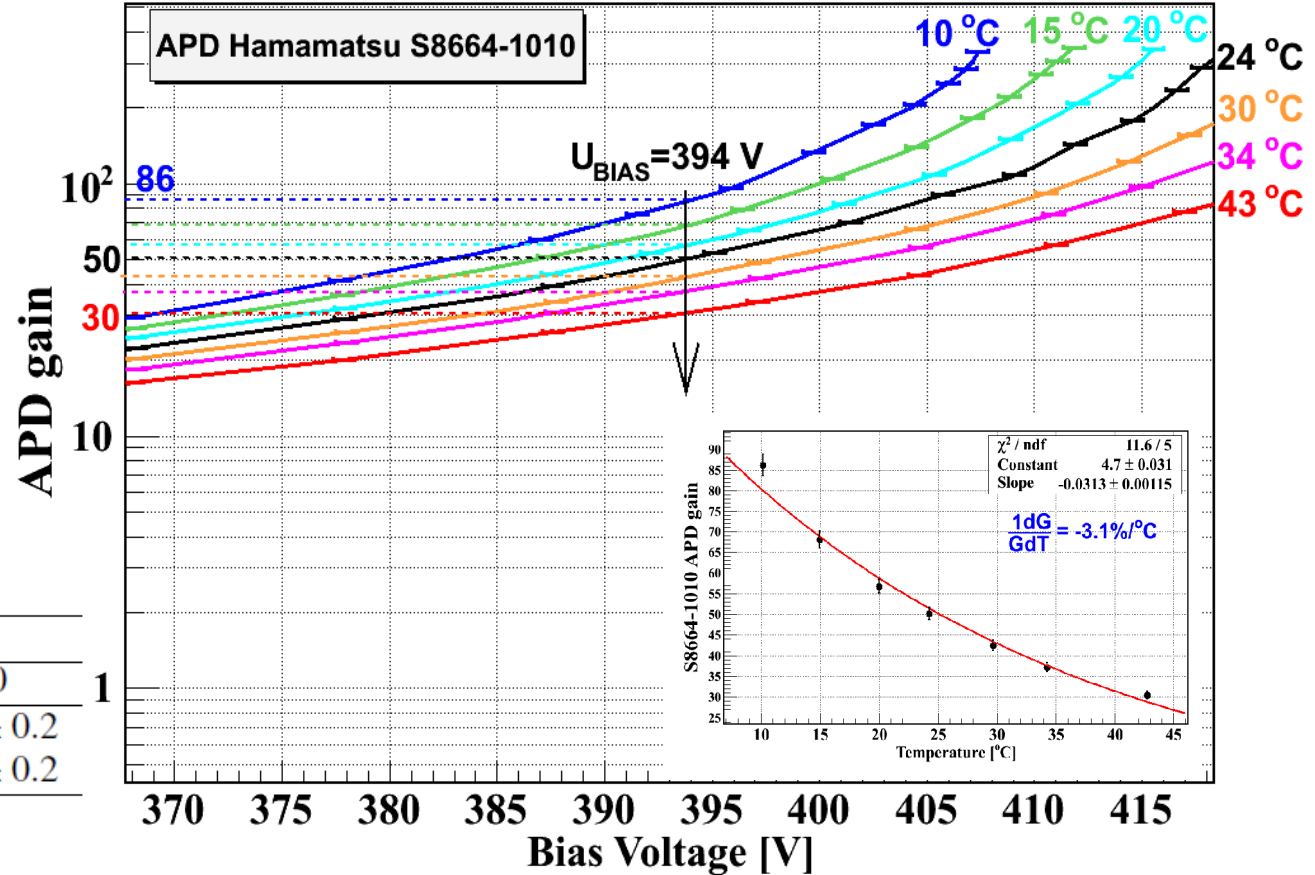
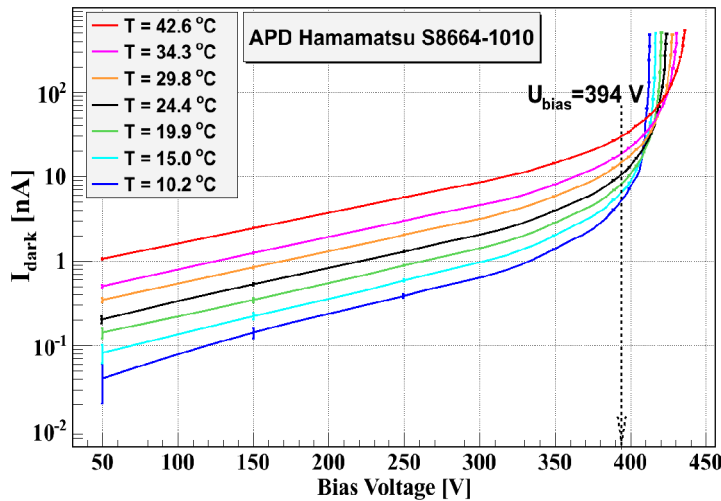


ENE(2 S8664-1010) = (0.55 ± 0.05) MeV
ENE(4 S8664-55) = (0.45 ± 0.05) MeV



Characteristics of APD

APD is a dominant source of the signal temperature variations, which have to be compensated



$(1/G)(dG/dT)[\%/^{\circ}\text{C}]$

APD type	APD gain		
	30	50	100
S8664-55	-2.4 ± 0.1	-3.3 ± 0.1	-5.0 ± 0.2
S8664-1010	-2.3 ± 0.1	-3.1 ± 0.1	-4.9 ± 0.2

To compensate temperature variations of APD gain, we can organize temperature sensor - bias voltage feedback.

Summary

- Hamamatsu APDs of S8664 series provide promising option for Belle II end cap ECL upgrade
- Essential increase of the light output of the CsI(pure)+APD(s) counter was achieved with WLS plates based on the nanostructured organosilicon luminophores (NOL-9)
- Several APDs per crystal allow us to decrease further ENE and provide readout redundancy
- The ENE of the counter with 4 S8664-55 APDs was measured to be $ENE = (0.45 \pm 0.05)$ MeV, which satisfies project requirements
- Radiation hardness of WLS plates is under investigation
- We are working to optimize the readout scheme from WLS plates